# A COVARIANT AND EXTENDED APPROACH TO <br> SOME PHYSICAL PROBLEMS WITH <br> CONSTRAINED FIELD VARIABLES 

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#### Abstract

Many physical problems are described by means of systems $S$ of partial differential equations, whose field variables $\underline{u}$ are restricted by relations of the type $\Phi_{I}(\underline{u})=0$. Some examples, to this regard, are the ultrarelativistic gases studied in the framework of Extended Thermodynamics, the relativistic magnetofluiddynamics and the Maxwell Equations in the relativistic form. Here a general method is proposed to deal with problems of this kind; in particular, a new system $S^{\prime}$ is proposed in the independent variables $\underline{u}, \psi_{R}$ which are not restricted. Moreover, the solutions of $S^{\prime}$ with $\Phi_{I}(\underline{u})=0, \psi_{R}=$ 0 , are the same of the original system $S$. The new system $S^{\prime}$ is expressed in the covariant form and is hyperbolic, under the assumption that the original system $S$ satisfies these properties; $\Phi_{I} \underline{(u)}=0, \psi_{R}=0$ are satisfied as consequences of $S^{\prime}$ and of the intial conditions. The new variables $\psi_{R}$ are only auxiliary quantities.


## 1. Introduction.

Let us consider the physical problems which are described by means of a quasi-linear system of partial differential equations of the type

$$
\begin{equation*}
\sum_{j=1}^{N} A_{i j}^{\alpha}(\underline{u}) \partial_{\alpha} u_{j}=f_{i}(\underline{u}) \tag{1}
\end{equation*}
$$

Entrato in Redazione il 24 marzo 1997.
for $i=1,2, \ldots, n$ and in the $N$ variables $u_{1}, \ldots, u_{N}$.
When the variables $\underline{u}$ are independent and $n=N$, this system is un-constrained, and its hyperbolicity is easily defined as follows.

Definition 1. The system (1) is hyperbolic ([3], [4]) in the time-direction $t^{\alpha}$ (such $t^{\alpha} t_{\alpha}=-1$ ) if and only if

1) $\operatorname{det}\left(A_{i j}^{\alpha} t_{\alpha}\right) \neq 0$;
2) for any four-vector $n^{\alpha}$ such that $n^{\alpha} t_{\alpha}=0, n^{\alpha} n_{\alpha}=1$, the eigenvalue problem $\sum_{j=1}^{N} A_{i j}^{\alpha}\left(n_{\alpha}-\lambda t_{\alpha}\right) \delta u_{j}=0$ has real eigenvalues $\lambda$ and $N$ linearly independent l.i eigenvectors $\delta u_{j}$.
We notice that the first of these conditions is equivalent to the following one
$1^{\prime}$ ) the system $\left.t_{\alpha} \sum_{j=1}^{N} A_{i j}^{\alpha} \underline{( }\right) \delta u_{j}=0$, in the independent unknowns $\delta u_{j}$, has only the solution $\delta u_{j}=0$.
However, in some physical problems, the $N$ variables $\underline{u}$ are not independent, but constrained by $M$ relations of the type

$$
\begin{equation*}
\Phi_{I}(\underline{u})=0 \quad \text { for } \quad I=1, \ldots, M \tag{2}
\end{equation*}
$$

where the functions $\Phi_{I}$ are differentiable with respect to $\underline{u}$, expressed in covariant form, and functionally independent, i.e. the rectangular matrix

$$
\frac{\partial \Phi_{I}}{\partial u_{j}} \quad I=1, \ldots, M ; j=1, \ldots, N
$$

has rank $M$. In this way there remain $N-M$ independent variables.
If $n=N-M$, the above definition can be easily extended and becomes
Definition 2. The system (1) under the constraints (2) is hyperbolic in the timedirection $t_{\alpha}$ if and only if

1) the system $t_{\alpha} \sum_{j=1}^{N} A_{i j}^{\alpha}(\underline{u}) \delta u_{j}=0, \sum_{j=1}^{N} \frac{\partial \Phi_{I}}{\partial u_{j}} \delta u_{j}=0$, in the independent unknowns $\delta u_{j}$, has only the solution $\delta u_{j}=0$;
2) the problem

$$
\left(n_{\alpha}-\lambda t_{\alpha}\right) \sum_{j=1}^{N} A_{i j}^{\alpha}(\underline{u}) \delta u_{j}=0, \quad \sum_{j=1}^{N} \frac{\partial \Phi_{I}}{\partial u_{j}} \delta u_{j}=0
$$

in the independent unknowns $\delta u_{j}$, has real eigenvalues $\lambda$ and $N$ 1.i eigenvectors $\delta u_{j}$.

In fact, if we choose $N-M$ parameters $q_{h}$ such that $u_{j}=u_{j}\left(q_{h}\right)$ is the general solution of the constraints (2), the Definition 2 is exactly the Definition 1 written in the $N-M$ un-constrained variables $q_{h}$. This approach has been widely used in literature; see, for example, ref. [7] where the independent variables are $n$ (partile density), $e$ (energy density), $\pi$ (dinamic pressure), $U^{\alpha}$ (4-velocity), $q^{\alpha}$ (heat flux), $t^{\langle\alpha \beta\rangle}$ (stress deviator), constrained by

$$
U^{\alpha} U_{\alpha}=-1, q^{\alpha} U_{\alpha}=0, t^{\langle\alpha \beta\rangle} U_{\alpha}=0, t^{\langle\alpha \beta\rangle} g_{\alpha \beta}=0
$$

with $g_{\alpha \beta}$ the metric tensor.
Finally, if $n>N-M$, we have also $n-(N-M)$ differential constraints. In this case we may conceive the idea of taking only $N-M$ equations from the system (1) and hoping that this reduced system is hyperbolic, according to Definition 2; more generally, we may take $N-M$ linear combinations of the equations of system (1), or, equivalently, we may multiply it on the left by a $(N-M) X n$ matrix $Z_{k i}$. In order not to lose manifest covariance, we may allow the matrix $Z_{k i}$ to have more than $N-M$ rows, but to have rank $N-M$; in this way the supplementary equations are only linear combinations of the others. Therefore, the following definition of hyperbolicity, for this constrained system, is proposed:

Definition 3. The system (1) under the constraints (2) is hyperbolic in the timedirection $t_{\alpha}$ if and only if an $m$ n matrix $Z_{k i}$ exists, such that,

1) $Z_{k i}$ has rank $N-M$;
2) the system $t_{\alpha} \sum_{i=1}^{n} \sum_{j=1}^{N} Z_{k i} A_{i j}^{\alpha}(\underline{u}) \delta u_{j}=0, \sum_{j=1}^{N} \frac{\partial \Phi_{I}}{\partial u_{j}} \delta u_{j}=0$, in the independent unknowns $\delta u_{j}$, has only the solutions $\delta u_{j}=0$;
3) the problem

$$
\left(n_{\alpha}-\lambda t_{\alpha}\right) \sum_{i=1}^{n} \sum_{j=1}^{N} Z_{k i} A_{i j}^{\alpha}(\underline{u}) \delta u_{j}=0, \sum_{j=1}^{N} \frac{\partial \Phi_{I}}{\partial u_{j}} \delta u_{j}=0
$$

in the independent unknowns $\delta u_{j}$, has real eigenvalues $\lambda$ and $N$ l.i eigenvectors $\delta u_{j}$.
If the system (1) is expressed in covariant form, also the matrix $Z_{k i}$ must be covariant, to preserve this property.
Obviously, we may substitute the system (1) with

$$
\begin{equation*}
\sum_{i=1}^{n} \sum_{j=1}^{N} Z_{k i} A_{i j}^{\alpha}(\underline{u}) \partial_{\alpha} u_{j}=\sum_{i=1}^{n} Z_{k i} f_{i}(\underline{u}), \tag{3}
\end{equation*}
$$

and apply to it the Definition 2.
I remark that this arguments are inspired by the elegant Strumia's papers on this subject ([10], [11], [12]), altough the definition in this article is less restrictive than Strumia's one, as it will be seen in section 2. Also in this section, a method will be shown to find the matrix $Z_{k i}$.
The above definition of hyperbolicity for a constrained system appears a little complicated. To eliminate this drawback I propose a method based on the ideas of extended thermodynamics ([6], [7]) to introduce other independent variables $\psi_{R}$ and to find a new system of equations that for $\psi_{R}=0$ reduces to (1); moreover this new system has the same number of equations and of independent variables, it is hyperbolic if and only if the system (1) is hyperbolic, it gives $\psi_{R}=0$ if we impose $\psi_{R}=0$ only on a given time-like initial hypersurface, it is expressed in covariant form. These results will be obtained in Section 3 and are expressed by the system (17).

Another idea is that of searching a new system with a less number of auxiliary variables $\psi_{R}$ than in (17), and without the constraints (2); this idea is realized in Section 4 and the new system is expressed by (24).
All these results depend on the time-like congruence $t_{\alpha}$ that has been initially chosen. This problem is investigated in Section 5, under the assumption that the system (1) is hyperbolic in the direction $t_{\alpha}$ and the characteristic velocities do not exceed the speed of light.
Physical examples of application of this methodology are also considered in this paper; they are the equations of Extended Thermodynamics for ultra-relativistic gases (see in Section 2), those of relativistic fluiddynamics (see in Section 4 from eq. (30) to eq. (32)), those of relativistic magnetofluiddynamics (see in Section 3 from eq. (21) and Section 4, eq. (29)), those of covariant Maxwell electro-dynamics (see in Section 4 from eq. (33)).

It is also shown, in Section 2 from eq. (13), that the Einstein's equations in empty space are not hyperbolic.

## 2. A method to find the matrix $Z_{k i}$.

If system (1) is hyperbolic, it can be written in the form (3) which makes more easy to find the associated eigenvalues and eigenvectors. The problem now arises on how to find the matrix $Z_{k i}$. To this end, the following 4 steps can be accomplished.
Step 1) Let us consider the system $\delta \Phi_{I}=0$, of $M$ linearly independent
equations in the $N$ unknowns $\delta u_{j}$, i.e.,

$$
\begin{equation*}
\sum_{j=1}^{N} \frac{\partial \Phi_{I}}{\partial u_{j}} \delta u_{j}=0 \tag{4}
\end{equation*}
$$

which is a consequence of the constraints (2); it gives $\delta u_{j}$ as a linear combination of $N-M$ free unknowns. Another possibility, useful in order not to lose the covariance, is to obtain $\delta u_{j}$ as a linear combination of $p$ free unknowns, with $p \geq N-M$, but by means of functions which are not functionally independent. More precisely, we may find

$$
\begin{equation*}
\delta u_{j}=U_{j j^{\prime}}(\underline{u}) V_{j^{\prime}}, \forall V_{j^{\prime}}, \tag{5}
\end{equation*}
$$

with $U_{j j^{\prime}}$ a matrix of rank $N-M$.
For the applications it is useful to notice that $U_{j j^{\prime}}$ is the matrix whose $j$-th row is the derivative of $\delta u_{j}$ in equation (5), with respect to $V_{j^{\prime}}$. This matrix is such that

$$
\begin{equation*}
\sum_{j=1}^{N} \frac{\partial \Phi_{I}}{\partial u_{j}} U_{j j^{\prime}}=0 \tag{6}
\end{equation*}
$$

The practical meaning of this step is that the variables contribute to the equations, only by means of their projections onto the subspace tangential to the variety (2).
Step 2) Let us consider the matrix $t_{\alpha} \sum_{j=1}^{N} A_{i j}^{\alpha} U_{j j^{\prime}} ;$ it has rank $N-M$, as proved in appendix 1 . For the applications, we notice that its $i$-th row is the derivative with respect to $V_{j^{\prime}}$ of the expression $t_{\alpha} \sum_{j=1}^{N} A_{i j}^{\alpha} \delta u_{j}$, after having substituted $\delta u_{j}$ from equation (5).
After that, let us consider the system with this matrix of the coefficients and contracted on the left by the unknowns $X_{i}$,

$$
\begin{equation*}
\sum_{i=1}^{n} X_{i} t_{\alpha} \sum_{j=1}^{N} A_{i j}^{\alpha} U_{j j^{\prime}}=0 \tag{7}
\end{equation*}
$$

I gives $X_{i}$ as a linear combination of $n-(N-M)$ free unknowns. Alternatively, we may find $X_{i}$ as a linear combination of $q$ free unknowns, with $q \geq$
$n-(N-M)$, but by means of functionally dependent functions; this possibility is used in order not to lose the covariance. More precisely, we may find

$$
\begin{equation*}
X_{i}=\sum_{i^{\prime}=1}^{q} \bar{X}_{i^{\prime}} X_{i^{\prime} i} \quad \forall \bar{X}_{i^{\prime}} \tag{8}
\end{equation*}
$$

where the $q X n$ matrix $X_{i^{\prime} i}$ has rank $n-(N-M)$.
(Note that $X_{i^{\prime} i}$ is the matrix whose i-th coulumn is the derivative of $X_{i}$ in equation (8) with respect to $\bar{X}_{i^{\prime}}$ ).
This matrix is such that the following relation holds

$$
\begin{equation*}
\sum_{i=1}^{n} X_{i^{\prime} i} t_{\alpha} \sum_{j=1}^{N} A_{i j}^{\alpha} U_{j j^{\prime}}=0 \tag{9}
\end{equation*}
$$

The practical meaning of this matrix is that it allows to separate the differential constraints from the other equations.
Step 3) Let us consider the system

$$
\begin{equation*}
\sum_{k=1}^{n} X_{i^{\prime} k} Y_{k}=0 \tag{10}
\end{equation*}
$$

it gives

$$
\begin{equation*}
Y_{k}=\sum_{i=1}^{n} Y_{k i} \bar{Y}_{i}, \quad \forall \bar{Y}_{i} \quad \text { (free unknowns) } \tag{11}
\end{equation*}
$$

where $Y_{k i}$ is a $n \times n$ matrix having rank $N-M$. Obviously, it is such that

$$
\begin{equation*}
\sum_{k=1}^{n} X_{i^{\prime} k} Y_{k i}=0 \tag{12}
\end{equation*}
$$

Here too, the $k$-th row of $Y_{k i}$ is the derivative of $Y_{k}$ in eq. (11) with respect to $\bar{Y}_{i}$. The practical meaning of multiplyng the original system on the left by $Y_{k i}$ is that of projecting it onto the subspace ortogonal to $X_{i^{\prime} i}$; in this way, the constant (9) on the evolutive part of the equations becomes now a constraint also on its spatial part.
Step 4) The matrix $Z_{k i}$ is sum of $Y_{k i}$ and of a suitable solution of the system (9), i.e., the parameters $\mu_{k i}$, exist such that

$$
Z_{k i}=Y_{k i}+\sum_{i^{\prime}=1}^{n} \mu_{k i}, X_{i^{\prime} i}
$$

Obviously, the parameters $\mu_{k i^{\prime}}$ must be such that the above mentioned properties of the matrix $Z_{k i}$, are preserved. The second term, in the expression of $Z_{k i}$, takes account of the fact that the differential constraints may still play a role, before to be neglected.
Having completed the description of these 4 step, let us prove the last of the them. The $2 n \times n$ matrix $\binom{Y_{i^{\prime} i}}{X_{i^{\prime} i}}$ has rank $n$ (see Appendix 2); therefore the parameters $\lambda_{k i^{\prime}}, \mu_{k i^{\prime}}$ exist such that $Z_{k i}=\sum_{i^{\prime}=1}^{n}\left(\lambda_{k i^{\prime}} Y_{i^{\prime} i}+\mu_{k i^{\prime}} X_{i^{\prime} i}\right)$; from this it follows

$$
t_{\alpha} \sum_{i=1}^{n} \sum_{j=1}^{N} Z_{k i} A_{i j}^{\alpha} \delta u_{j}=t_{\alpha} \sum_{i^{\prime}=1}^{n} \sum_{i=1}^{n} \sum_{j=1}^{N} \lambda_{k i^{\prime}} Y_{i^{\prime} i} A_{i j}^{\alpha} \delta u_{j}
$$

where (9) has been used and also the fact that $\delta u_{j}$ is a linear combination of $U_{j j^{\prime}}$. If we impose now the condition 2) of Definition 3, we obtain that $t_{\alpha} \sum_{i^{\prime}=1}^{n} \sum_{i=1}^{n} \sum_{j=1}^{N} \lambda_{k i^{\prime}} Y_{i^{\prime} i} A_{i j}^{\alpha} \frac{\partial u_{j}}{\partial q_{h}} \delta q_{h}=0$ has only the solution $\delta q_{h}=0$; this fact proves that the matrix $t_{\alpha} \sum_{i^{\prime}=1}^{n} \sum_{i=1}^{n} \sum_{j=1}^{N} \lambda_{k i^{\prime}} Y_{i^{\prime} i} A_{i j}^{\alpha} \frac{\partial u_{j}}{\partial q_{h}}$ has rank $N-M$ and, consequently, $\sum_{i^{\prime}=1}^{n} \lambda_{k i^{\prime}} Y_{i^{\prime} i}$ has rank $N-M$ (see Appendix 3).
In this way we see that $\sum_{i^{\prime}=1}^{n} \lambda_{k i^{\prime}} Y_{i^{\prime} i}$ is also a solution of system (12) and has rank $N-M$, i.e., $\sum_{i^{\prime}=1}^{n} \lambda_{k i^{\prime}} Y_{i^{\prime} i}$ satisfies the same properties of $Y_{k i}$; by substituting $\sum_{i^{\prime}=1}^{n} \lambda_{k i^{\prime}} Y_{i^{\prime} i}$ with $Y_{k i}$ we obtain $Z_{k i}=Y_{k i}+\sum_{i^{\prime}=1}^{n} \mu_{k i^{\prime}} X_{i^{\prime} i}$.
Therefore $Z_{k i}$ is determined except for $\mu_{k i^{\prime}}$; from equation (8), we see that $Z_{k i}$ is the sum of $Y_{k i}$ and of a particular solution of the system $\sum_{i=1}^{n} X_{i} t_{\alpha} \sum_{j=1}^{N} A_{i j}^{\alpha} U_{j j^{\prime}}=0$.
Let us introduce now some notes.
Note 1: We notice that the Definition 2 is less restrictive that Strumia's one [10]; in fact he considers only the case $n=N$ and imposes the further condition that the matrix $Z_{k i}$ is the projector onto the subspace generated by the 1.i. columns of $U_{j j^{\prime}}$.

Note 2: If system (1) is hyperbolic and $\delta u_{j}$ is a solution of the system $t_{\alpha} \sum_{j=1}^{N} A_{i j}^{\alpha} \underline{(\underline{u})} \delta u_{j}=0, \sum_{j=1}^{N} \frac{\partial \Phi_{I}}{\partial u_{j}} \delta u_{j}=0$, then $\delta u_{j}$ satisfies also the system
considered in condition 2) of Definition 2; therefore we have $\delta u_{j}=0$.
This note can be used to show that Einstein's equations in empty space are not hyperbolic. These equations are ([10], [2], [5]):

$$
\begin{equation*}
\partial_{\alpha} \Gamma_{\mu \nu}^{\alpha}-\partial_{\mu} \Gamma_{\nu \lambda}^{\lambda}+\Gamma_{\mu \nu}^{\lambda} \Gamma_{\lambda \rho}^{\rho}-\Gamma_{\mu \lambda}^{\rho} \Gamma_{\nu \rho}^{\lambda}=0 \tag{13}
\end{equation*}
$$

with $\Gamma_{\mu \nu}^{\alpha}=g^{\alpha \rho}\left(\partial_{\nu} g_{\rho \mu}+\partial_{\mu} g_{\rho \nu}-\partial_{\rho} g_{\mu \nu}\right) / 2$ (Christoffel symbol).
They constitute a second order system of ten equations for the unknown symmetric tensor $g_{\rho \mu}$.
By defining $\omega_{\alpha \mu \nu}=\partial_{\alpha} g_{\mu \nu}$, it can be reduced to the first order system

$$
\left\{\begin{array}{l}
\partial_{\alpha}\left[g^{\alpha \rho}\left(\omega_{\nu \rho \mu}+\omega_{\mu \rho \nu}-\omega_{\rho \mu \nu}\right)-\delta_{\mu}^{\alpha} g^{\lambda \rho} \omega_{\nu \rho \lambda}\right]=W_{\mu \nu}\left(g_{\alpha \delta}, \omega_{\beta \gamma \delta}\right)  \tag{14}\\
\partial_{\alpha}\left(\delta_{\beta}^{\alpha} g_{\mu \nu}\right)=\omega_{\beta \mu \nu} \\
\left(\delta_{\sigma}^{\alpha} \delta_{\tau}^{\beta}-\delta_{\tau}^{\alpha} \delta_{\sigma}^{\beta}\right) \partial_{\alpha} \omega_{\beta \mu \nu}=0
\end{array}\right.
$$

in the 50 unknowns $g_{\mu \nu}, \omega_{\beta \mu \nu}$.
We have $I=0$ and $U_{j j^{\prime}}=\delta_{j j^{\prime}}$, because there is no constraint on the independent variables. However, the condition on note 3 , which is necessary for the hyperbolicity of the system (14) is not satisfied. In fact, in this case, the system in note 2 is

$$
\left\{\begin{array}{l}
t_{\alpha} \delta\left[g^{\alpha \rho}\left(\omega_{\nu \rho \mu}+\omega_{\mu \rho \nu}-\omega_{\rho \mu \nu}\right)-\delta_{\mu}^{\alpha} g^{\lambda \rho} \omega_{\nu \rho \lambda}\right]=0  \tag{15}\\
t_{\alpha}\left(\delta_{\beta}^{\alpha} g_{\mu \nu}\right)=0 \\
\left(\delta_{\sigma}^{\alpha} \delta_{\tau}^{\beta}-\delta_{\tau}^{\alpha} \delta_{\sigma}^{\beta}\right) t_{\alpha} \delta \omega_{\beta \mu \nu}=0
\end{array}\right.
$$

Its general solution is

$$
\delta g_{\mu \nu}=0, \delta \omega_{\tau \mu \nu}=t_{\tau}\left(t_{\mu} V_{\nu}+t_{\nu} V_{\mu}+t_{\mu} t_{\nu} t^{\lambda} V_{\lambda}\right)
$$

with $V_{\mu}$ an arbitrary four-vector. Therefore this system has solutions different from $\delta u_{j}=0$, i.e., the Einstein's equations in empty space are not hyperbolic. I conclude this section by illustrating the above 4 steps in the case of Extended Thermodynamic of an ultrarelativistic gas. This problem is described by the equations

$$
\left\{\begin{array}{l}
\partial_{\alpha} V^{\alpha}=0 \\
\partial_{\alpha} T^{\alpha \beta}=0 \\
\partial_{\alpha} A^{\alpha \beta \gamma}=I^{\beta \gamma}
\end{array}\right.
$$

where the independent variables are $V^{\mu}$ (particle number-particle flux vector), $T^{\lambda \nu}$ (stress-energy-momentum tensor), constrained by $\Phi=0$, with $\Phi=T_{\lambda}^{\lambda}$.

Moreover, in the case of a non degenerate gas, we have

$$
\begin{gathered}
A^{\alpha \beta \gamma}=(4 / 3) n^{-1} e^{2}\left(2 U^{\alpha} U^{\beta} U^{\gamma}+g^{(\alpha \beta} U^{\gamma)}\right)+ \\
+2(e / n)\left(g^{(\alpha \beta} q^{\gamma)}+6 U^{(\alpha} U^{\beta} q^{\gamma)}\right)+6(e / n) t^{(\langle\alpha \beta\rangle} U^{\gamma)}, \\
I^{\beta \gamma}=B_{3} t^{\langle\beta \gamma\rangle}+2 B_{4} q^{(\beta} U^{\gamma)}, \\
n=\left(-V_{\beta} V^{\beta}\right)^{1 / 2}, U^{\alpha}=n^{-1} V^{\alpha}, \\
h^{\alpha \beta}=g^{\alpha \beta}+U^{\alpha} U^{\beta}, t^{\langle\alpha \beta\rangle}=\left(h_{\mu}^{\alpha} h_{v}^{\beta}-1 / 3 h^{\alpha \beta} h_{\mu \nu}\right) T^{\mu \nu} \\
q^{\alpha}=-h_{\mu}^{\alpha} U_{\nu} T^{\mu \nu}, e=U_{\mu} U_{\nu} T^{\mu \nu}
\end{gathered}
$$

An arbitrary single variable function $A$, which appears in the first version of these equations, has been dropped because it does not appear in other versions based on the kinetic theory.
For the sake of simplicity, let us take $t^{\alpha}=U^{\alpha}$.
Step 1) The system (4) is $g^{\alpha \beta} \delta T_{\alpha \beta}=0$, whose solution is $\delta T^{\lambda v}=\left(g_{\varphi}^{(\lambda} g_{\rho}^{\nu)}-\right.$ $\left.1 / 4 g^{\lambda v} g_{\varphi \rho}\right) W^{\varphi \rho}$; therefore, the requested matrix $U_{j j^{\prime}}$ is

$$
U_{j j^{\prime}}=\left(\begin{array}{cc}
g_{\delta}^{\mu} & 0 \\
0 & g^{(\lambda}{ }_{\varphi} g_{\rho}^{\nu)}-1 / 4 g^{\lambda v} g_{\varphi \rho}
\end{array}\right), \text { whose rank is } 9 .
$$

(The index $j$ is described by $\mu \lambda \nu$, while $j^{\prime}$ is described by $\delta \varphi \rho$ ).
Step 2) For the sake of simplicity, let us aim to hyperbolicity holding in a neighbourhood of thermodinamical equilibrium; in this case we may calculate the coefficients of $\partial_{\alpha} V^{\mu}, \partial_{\alpha} T^{\lambda v}$ at equilibrium. The resulting matrix $U_{\alpha} \sum_{j=1}^{N} A_{i j}^{\alpha} U_{j j^{\prime}}$ is

$$
\left(\begin{array}{cc}
U_{\delta} & 0 \\
0 & U_{(\varphi} g_{\rho)}^{\beta}-1 / 4 U^{\beta} g_{\varphi \rho} \\
A_{\delta}^{\beta \gamma} & A_{\varphi \rho}^{B \gamma}
\end{array}\right),
$$

with $\left.A_{\delta}^{\beta \gamma}=-(4 / 9) e^{2} n^{-2}\left[\left(h^{\beta \gamma}+3 U^{\beta} U^{\gamma}\right) U_{\delta}-10 h^{(\beta}{ }_{\delta} U^{\gamma}\right)\right]$,

$$
\begin{gathered}
A_{\varphi \rho}^{\beta \gamma}=(2 / 3)(e / n)\left[-4\left(g^{\beta \gamma}+4 / 3 U^{\beta} U^{\gamma}\right)\left(U_{\varphi} U_{\rho}+1 / 4 g_{\varphi \rho}\right)+\right. \\
\left.+10 U^{(\beta} h_{(\varphi}^{\gamma)} U_{\rho)}+h_{\varphi \rho} h^{\beta \gamma}-3 h_{(\varphi}^{\beta} h_{\rho)}^{\gamma}\right] .
\end{gathered}
$$

The system (7) is

$$
\begin{gathered}
X U_{\delta}+X_{\beta \gamma}(8 / 9) e^{2} n^{-2}\left[-2 U^{\beta} U^{\gamma} U_{\delta}+5 h_{\delta}^{\beta} U^{\gamma}\right]=0, \\
X_{\beta}\left[\left(U_{(\varphi} g_{\rho)}^{\beta}-1 / 4 U^{\beta} g_{\varphi \rho}\right]+X_{\beta \gamma}(2 / 9) e n^{-1}\left[-16 U^{\beta} U^{\gamma}\left(U_{\varphi} U_{\rho}+\right.\right.\right. \\
\left.\left.+1 / 4 g_{\varphi \rho}\right)+3\left(h_{\varphi \rho} U^{\beta} U^{\gamma}-3 h_{\varphi}^{\gamma} h_{\rho}^{\beta}\right)\right]=0,
\end{gathered}
$$

where the relation $h_{\delta}^{\beta} X_{\beta \gamma} U^{\gamma}=0$ (which comes from the first of these equations, contracted with $h_{\delta^{\prime}}^{\delta}$ ) has been used. Moreover, $X_{\beta \gamma}$ is a symmetric and traceless tensor.
The above system has the solution

$$
X=(16 / 3) e^{2} n^{-2} \bar{X}, X_{\beta}=(32 / 3) e n^{-1} \bar{X} U_{\beta}
$$

$X_{\beta \gamma}=\bar{X}\left(g_{\beta \gamma}+4 U_{\beta} U_{\gamma}\right)$. Consequently, the matrix $X_{i i^{\prime}}$ is $\left(i^{\prime}=1\right)$

$$
\left((16 / 3) e^{2} n^{-2},(32 / 3) e n^{-1} U_{\beta},\left(g_{\beta \gamma}+4 U_{\beta} U_{\gamma}\right)\right)
$$

Step 3) We have to consider now the system

$$
(16 / 3) e^{2} n^{-2} Y+(32 / 3) e n^{-1} U_{\beta} Y^{\beta}+\left(g_{\beta \gamma}+4 U_{\beta} U_{\gamma}\right) Y^{\beta \gamma}=0, g_{\beta \gamma} Y^{\beta \gamma}=0
$$

To find a covariant solution to this system, we notice that $Y_{1}, Y_{2}, \bar{Y}^{\lambda v}$ exist such that

$$
\begin{gathered}
Y^{\beta \gamma}=Y_{1} g^{\beta \gamma}+Y_{2} U^{\beta} U^{\gamma}+(1 / 3)\left[3 g_{\lambda}^{\beta} g_{\nu}^{\gamma}-\left(g_{\lambda \nu}+U_{\lambda} U_{\nu}\right) g^{\beta \gamma}+\right. \\
\left.-\left(4 U_{\lambda} U_{\nu}+g_{\lambda \nu}\right) U^{\beta} U^{\gamma}\right] \bar{Y}^{\lambda \nu}
\end{gathered}
$$

(In fact $Y_{1}, Y_{2}$ can be found by contracting this expression with $g_{\beta \gamma}+$ $U_{\beta} U_{\gamma},(1 / 3)\left(g_{\beta \gamma}+4 U_{\beta} U_{\gamma}\right)$ respectively; after that $\bar{Y}^{\lambda \nu}=Y^{\beta \gamma}$ is a possible choice for $\bar{Y}^{\lambda \nu}$ ). By using this expression the above equations become conditions which give $Y, Y^{\beta}, Y_{1}, Y_{2}$; the result is

$$
\begin{gathered}
Y=\bar{Y}, Y^{\beta}=g_{\lambda}^{\beta} \bar{Y}^{\lambda} \\
Y^{\beta \gamma}=-(4 / 9) e n^{-2}\left(e \bar{Y}+2 n U_{\lambda} \bar{Y}^{\lambda}\right)\left(g^{\beta \gamma}+4 U^{\beta} U^{\gamma}\right)+ \\
+(1 / 3)\left[3 g_{\lambda}^{\beta} g_{v}^{\gamma}-\left(g_{\lambda \nu}+U_{\lambda} U_{\nu}\right) g^{\beta \gamma}-\left(4 U_{\lambda} U_{v}+g_{\lambda \nu}\right) U^{\beta} U^{\gamma}\right] \bar{Y}^{\lambda \nu}
\end{gathered}
$$

after that, we see that the matrix $Y_{k i}$ is

$$
\left(\begin{array}{ccc}
1 & 0 & 0 \\
0 & g_{\beta}^{\beta^{\prime}} & 0 \\
e C^{\beta^{\prime} \gamma^{\prime}} & 2 n C^{\beta^{\prime} \gamma^{\prime}} U_{\beta} & C_{\beta \gamma}^{\beta^{\prime} \gamma^{\prime}}
\end{array}\right)
$$

where $C^{\beta^{\prime} \gamma^{\prime}}=-(4 / 9) e n^{-2}\left(4 U^{\beta^{\prime}} U^{\gamma^{\prime}}+g^{\beta^{\prime} \gamma^{\prime}}\right)$,

$$
\begin{aligned}
C_{\beta \gamma}^{\beta^{\prime} \gamma^{\prime}} & =g_{\beta}^{\beta^{\prime}} g_{\gamma}^{\gamma^{\prime}}-(1 / 3) g_{\beta \gamma} g^{\beta^{\prime} \gamma^{\prime}}-(1 / 3) U_{\beta} U_{\gamma} g^{\beta^{\prime} \gamma^{\prime}}+ \\
& -(1 / 3) U^{\beta^{\prime}} U^{\gamma^{\prime}} g_{\beta \gamma}-(4 / 3) U_{\beta} U_{\gamma} U^{\beta^{\prime}} U^{\gamma^{\prime}}
\end{aligned}
$$

Step 4) If we take $\mu_{k i^{\prime}}=0$, we obtain $Z_{k i}=Y_{k i}$ and the Definition 2 is satisfied, as it can be easily seen. For the sake of brevity, I avoid to report the proof of this statement. I simply note that the system multiplied on the left by $Z_{k i}$ give 3 equations; if we add to the third one of them the first one, premultiplied by $e C^{\beta^{\prime} \gamma^{\prime}}$, and the second one, premultiplied by $2 n C^{\beta^{\prime} \gamma^{\prime}} U_{\beta}$, we obtain

$$
\left\{\begin{array}{l}
\partial_{\alpha} V^{\alpha}=0 \\
\partial_{\alpha} T^{\alpha \beta}=0 \\
\partial_{\alpha} A^{\alpha \beta^{\prime} \gamma^{\prime}}-(1 / 3)\left(g^{\beta^{\prime} \gamma^{\prime}}+4 U^{\beta^{\prime}} U^{\gamma^{\prime}}\right) U_{\beta} U_{\gamma} \partial_{\alpha} A^{\alpha \beta \gamma}=I^{\beta^{\prime} \gamma^{\prime}}
\end{array}\right.
$$

where the identities $g_{\beta \gamma} A^{\alpha \beta \gamma}=0, U_{\beta} U_{\gamma} I^{\beta \gamma}=0, g_{\beta \gamma} I^{\beta \gamma}=0$, have been used. This system is the conterpart of (3), for this particular case.

## 3. An extended approach to systems with constrained field.

In this section a new system is searched, in the variables $\underline{u}, \Psi_{R}$ (auxiliary quantities), which for $\Psi_{R}=0$ reduces to system (1); if (1) is hyperbolic, then also the new system is hyperbolic.
To this end, let $T_{k}^{R}$ be $n-(N-M)$ 1.i. solutions of the system

$$
\begin{equation*}
T_{k} Z_{k i}=0 \tag{16}
\end{equation*}
$$

expressed in covariant form and let $\omega^{\alpha}$ be a four-vectorial function such that $t^{\alpha} \omega_{\alpha} \neq 0$ (for example we may choose $\omega^{\alpha}=t^{\alpha}$ ).
Let us consider the system

$$
\begin{equation*}
\sum_{i=1}^{n} \sum_{j=1}^{N} Z_{k i} A_{i j}^{\alpha}(\underline{u}) \partial_{\alpha} u_{j}+\sum_{R=1}^{n-(N-M)} \omega^{\alpha} T_{k}^{R} \partial_{\alpha} \Psi_{R}=\sum_{i=1}^{n} Z_{k i} f_{i}(\underline{u}) \tag{17}
\end{equation*}
$$

in the variables $u_{j}, \Psi_{R}$ constrained by $\Phi(\underline{u})=0$, so that we have $N-M+$ $[n-(N-M)]=n$ independent variables.
This system has the advantage to have an equal number of equations and of independent variables; when $\Psi_{R}=0$ it reduces to system (10), which is equivalent to system (1) except for some conditions to be imposed in the initial manyfold (see the previous section).
Moreover we prove that

1) the system (17) is hyperbolic if this property is satisfied by the system (1),
2) if $\Psi_{R}=0$ on an intial hypersurface $\Sigma$, then $\Psi_{R}=0$ will propagate off $\Sigma$.

In fact the system (17) is equivalent to

$$
\begin{equation*}
\sum_{i=1}^{n} \sum_{j=1}^{N} Z_{k i} A_{i j}^{\alpha} \frac{\partial u_{j}}{\partial q_{h}} \partial_{\alpha} q_{h}+\sum_{R=1}^{n-(N-M)} \omega^{\alpha} T_{k}^{R} \partial_{\alpha} \psi_{R}=\sum_{i=1}^{n} Z_{k i} f_{i} \tag{18}
\end{equation*}
$$

Condition 1) of Definition 1 is equivalent to impose that the system

$$
\begin{equation*}
t_{\alpha} \sum_{i=1}^{n} \sum_{j=1}^{N} Z_{k i} A_{i j}^{\alpha} \frac{\partial u_{j}}{\partial q_{h}} \delta q_{h}+\sum_{R=1}^{n-(N-M)} t_{\alpha} \omega^{\alpha} T_{k}^{R} \delta \psi_{R}=0 \tag{19}
\end{equation*}
$$

has only the solution $\delta q_{h}=\delta \psi_{R}=0$.
If we multiply eq. (19) by $T_{k}^{S}$, we obtain $\sum_{R=1}^{n-(N-M)} t_{\alpha} \omega^{\alpha} T_{k}^{S} T_{k}^{R} \delta \psi_{R}=0$, from which $\delta \psi_{R}=0$ because $T_{k}^{R}$ are 1.i.; after that, the system (19) gives $\delta q_{h}=0$ because the system (1) is hyperbolic and condition 2) of Definition 3 is satisfied. Let us now see that the system (18) satisfies also the condition 2) of Definition 1. To this end, let $\delta q_{k}$, for $k=1, \ldots, N-M$, be $N-M$ l.i. solutions of the condition 2) in Definition 2, corresponding to the eigenvalue $\lambda_{k}$; let $\delta \psi_{R}$ be the vector of components $\delta \psi_{R P}=\delta_{R P}$.
The vectors $\left(\delta q_{k}, 0_{R}\right)$ and $\left(0_{k}, \delta \psi_{P}\right)$ are $n$ 1.i. eigenvectors of the system (18), corresponding to the eigenvalues $\lambda_{k}$ and $\lambda^{\prime}=\left(n_{\alpha} \omega^{\alpha}\right) /\left(t_{\alpha} \omega^{\alpha}\right)$, respectively. In fact they satisfy the system

$$
\left(n_{\alpha}-\lambda t_{\alpha}\right)\left[\sum_{i=1}^{n} \sum_{j=1}^{N} Z_{k i} A_{i j}^{\alpha} \frac{\partial u_{j}}{\partial q_{h}} \delta q_{h}+\sum_{R=1}^{n-(N-M)} \omega^{\alpha} T_{k}^{R} \delta \psi_{R}\right]=0
$$

In this way we have proved that the system (17) is hyperbolic. Let us prove now the statement 2). To this end, we multiply the system (17) by $T_{k}^{S}$, obtaining that

$$
\sum_{R=1}^{n-(N-M)} T_{k}^{S} T_{k}^{R} \omega^{\alpha} \partial_{\alpha} \psi_{R}=0, \quad \text { i.e., } \quad \omega^{\alpha} \partial_{\alpha} \psi_{R}=0
$$

this relation can be written also as

$$
\begin{equation*}
t^{\alpha} \partial_{\alpha} \psi_{R}=\left(t^{\mu} \omega_{\mu}\right)^{-1} \omega_{\gamma} h^{\gamma \alpha} \partial_{\alpha} \psi_{R} \tag{20}
\end{equation*}
$$

with $h^{\gamma \alpha}=g^{\gamma \alpha}+t^{\gamma} t^{\alpha}$.

If $\psi_{R}=0$ on the initial hypersurface $\Sigma$ (from which $h^{\gamma \alpha} \partial_{\alpha} \psi_{R}=0$ also holds in $\Sigma$ ), we see from eq. (20) that $\psi_{R}=0$ holds also off $\Sigma$, as we desired to prove.
Let us now apply this methdology to a particular physical problem, the equations of relativistic magnetofluiddynamics. These equations are the conservation laws of mass and energy-momentum and the maxwell's equations, i.e.,

$$
\left\{\begin{array}{l}
\partial_{\alpha}\left(n u^{\alpha}\right)=0 \\
\partial_{\alpha}\left[\left(e+p+b^{2}\right) u^{\alpha} u^{\beta}+\left(p+b^{2} / 2\right) g^{\alpha \beta}-b^{\alpha} b^{\beta}\right]=0 \\
\partial_{\alpha}\left(u^{\alpha} b^{\beta}-b^{\alpha} u^{\beta}\right)=0
\end{array}\right.
$$

in the variables e (total energy-density), $V^{\alpha}, b^{\alpha}$ (related to the electromagnetic field), constrained by $V_{\alpha} b^{\alpha}=0$. Moreover, we have $n=\left(-V^{\gamma} V_{\gamma}\right)^{-1 / 2}$ (restmass density), $u^{\alpha}=n^{-1} V^{\alpha}$ (four-velocity), $p=p(n, e)$.
Hyperbolicity of this system has already been proved, for example in refs. ([10],
[11],[9], [1], [8]). By appliying the present methodology, with $t^{\alpha}=u^{\alpha}, \omega^{\alpha}=$ $u^{\alpha}$, we obtain that $n=9, N=9, M=1$,

$$
\begin{gathered}
U_{j j^{\prime}}=\left(\begin{array}{ccc}
1 & 0_{\gamma} & 0_{\gamma} \\
0_{\gamma} & g_{\gamma \delta} & 0_{\gamma \delta} \\
0_{\gamma} & n^{-1} b_{\delta} u_{\gamma} & h_{\gamma \delta}
\end{array}\right) ; X_{i^{\prime} i}=\left(\begin{array}{ccc}
0 & 0_{\beta} & u_{\beta} \\
0_{\gamma} & 0_{\gamma \beta} & 0_{\gamma \beta} \\
0_{\gamma} & 0_{\gamma \beta} & 0_{\gamma \beta}
\end{array}\right) ; \\
Y_{k j}=\left(\begin{array}{ccc}
1 & 0^{\beta} & 0^{\beta} \\
0^{\lambda} & g^{\lambda \beta} & 0^{\lambda \beta} \\
0^{\lambda} & 0^{\lambda \beta} & h^{\lambda \beta}
\end{array}\right) ; \mu_{k i^{\prime}}=\left(\begin{array}{ccc}
0 & 0^{\gamma} & 0^{\gamma} \\
b^{\lambda} & 0^{\lambda \gamma} & 0^{\lambda \gamma} \\
0^{\lambda} & 0^{\lambda \gamma} & 0^{\lambda \gamma}
\end{array}\right) ; \\
Z_{k j}=\left(\begin{array}{ccc}
1 & 0^{\beta} & 0^{\beta} \\
0^{\lambda} & g^{\lambda \beta} & b^{\lambda} u^{\beta} \\
0^{\lambda} & 0^{\lambda \beta} & h^{\lambda \beta}
\end{array}\right) ; T_{k}^{1}=\left(\begin{array}{c}
1 \\
0^{\lambda} \\
u^{\lambda}
\end{array}\right) .
\end{gathered}
$$

Consequently, the system (17), in this case, is

$$
\left\{\begin{array}{l}
\partial_{\alpha}\left(n u^{\alpha}\right)=0  \tag{22}\\
\partial_{\alpha}\left[\left(e+p+b^{2}\right) u^{\alpha} u^{\lambda}+\left(p+b^{2} / 2\right) g^{\alpha \lambda}-\right. \\
\left.\quad-b^{\alpha} b^{\lambda}\right]+u^{\gamma} b^{\lambda} \partial_{\alpha}\left(u^{\alpha} b_{\gamma}-b^{\alpha} u_{\gamma}\right)=0 \\
h^{\beta \gamma} \partial_{\alpha}\left(u^{\alpha} b_{\beta}-b^{\alpha} u_{\beta}\right)+u^{\lambda} u^{\alpha} \partial_{\alpha} \psi=0
\end{array}\right.
$$

Obviously, the system (22) in the unknowns $e, V^{\alpha}, b^{\alpha}, \psi$ constrained by $V_{\alpha} b^{\alpha}=0$ is hyperbolic, is expressed in covariant form and has an equal number of equations and of independent variables.

## 4. A new form for the system (17).

Let us consider firstly the case $n \geq N$ (from which $n-(N-M) \geq M$ holds), which I call case $a$. The system (17) in the independent variables $q_{1}, \ldots, q_{N-M}, \psi_{1}, \ldots, \psi_{M}, \psi_{M+1}, \ldots, \psi_{n-(N-M)}$ is hyperbolic. There are $M$ columns of the matrix $\frac{\partial \Phi_{I}}{\partial u_{j}}$, which are 1.i.; we can suppose, without loss of generality, that they are the last $M$ columns, i.e., that the matrix

$$
\frac{\partial \Phi_{I}}{\partial u_{j}} \text { for } j=N-M+1, \ldots, N \text { has rank } M
$$

Therefore we can take $q_{1}=u_{1}, \ldots, q_{N-M}=u_{N-M}$ and the system (17) in the independent variables $u_{1}, \ldots, u_{N-M}, \psi_{1}, \ldots, \psi_{M}, \psi_{M+1}, \ldots \psi_{n-(N-M)}$ is hyperbolic. Let us consider the invertible change of variables, expressing the above ones in terms of $u_{1}, \ldots, u_{N-M}, u_{N-M+1}, \ldots, u_{N}, \psi_{M+1}, \ldots, \psi_{n-(N-M)}$, by means of the transformation law

$$
\begin{equation*}
\Phi_{I}(\underline{u})=\psi_{I} \quad \text { for } I=1,2, \ldots, M . \tag{23}
\end{equation*}
$$

In other words, we are obtaining $u_{N-M+1}, \ldots, u_{N}$ as functions of the previous variables, from eq. (23).
Obviously, the system (17), expressed in terms of the new variables, is still hyperbolic. By using eq. (23) it becomes

$$
\begin{align*}
& \sum_{i=1}^{n} \sum_{j=1}^{N} Z_{k i} A_{i j}^{\alpha} \partial_{\alpha} u_{j}+\sum_{R=1}^{M} \omega^{\alpha} T_{k}^{R} \partial_{\alpha} \Phi_{R}(\underline{u})+  \tag{24}\\
& \quad+\sum_{R=M+1}^{n-(N-M)} \omega^{\alpha} T_{k}^{R} \partial_{\alpha} \psi_{R}=\sum_{i=1}^{n} Z_{k i} f_{i} .
\end{align*}
$$

This new form for the system (17) has the advantage to have an equal number of equations and of independent variables; these last ones are $\underline{u}, \psi_{M+1}, \ldots$, $\psi_{n-(N-M)}$ and are not constrained.
However, if we impose that $\Phi_{I}(\underline{u})=0, \psi_{M+1}=0, \ldots, \psi_{n-(N-M)}=0$ hold on the initial manifold, they will be satisfied also off it.
Let us consider now the case $n<N$ (from which $n-(N-M)<M$ holds), which I call case $b$.
There are $M$ columns of the matrix $\frac{\partial \Phi_{I}}{\partial u_{j}}$, which are 1.i.; we can suppose, without loss of generality, that they are the last $M$ columns, i.e., that the matrix
$M^{*}=\frac{\partial \Phi_{I}}{\partial u_{j}}$ for $j=N-M+1, \ldots, N$, has rank $M$. If we use the Laplace's rule to calculate $\operatorname{det} M^{*} \neq 0$, we see that we can exchange some columns of $M^{*}$ such that $M^{*}$ becomes $\left(\begin{array}{cc}M_{11}^{*} & M_{12}^{*} \\ M_{21}^{*} & M^{*} 22\end{array}\right)$ with $M_{11}^{*}$ of order $n-(N-M)$, $M_{22}^{*}$ of order $N-n$ and moreover $\operatorname{det} M_{11}^{*} \neq 0$, $\operatorname{det} M_{22}^{*} \neq 0$. In other words, by exchanging the names of some $u_{j}$, we obrtain that $\frac{\partial \Phi_{I}}{\partial u_{j}}$ takes the form

$$
\left(\begin{array}{ccc}
\frac{\partial \Phi_{I^{\prime}}}{\partial u_{j^{\prime}}} & \frac{\partial \Phi_{I^{\prime}}}{\partial u_{j^{\prime \prime}}} & \frac{\partial \Phi_{I^{\prime}}}{\partial u_{j^{\prime \prime \prime}}} \\
\frac{\partial \Phi_{I^{\prime \prime}}}{\partial u_{j^{\prime}}} & \frac{\partial \Phi_{I^{\prime \prime}}}{\partial u_{j^{\prime \prime}}} & \frac{\partial \Phi_{I^{\prime \prime \prime}}}{\partial u_{j^{\prime \prime \prime}}}
\end{array}\right)
$$

for $I^{\prime}=1, \ldots, n-(N-M) ; I^{\prime \prime}=n-(N-M)+1, \ldots, M$;

$$
j^{\prime}=1, \ldots, N-M ; \quad j^{\prime \prime}=N-M+1, \ldots, n ; \quad j^{\prime \prime \prime}=n+1, \ldots, N
$$

and we have

$$
\begin{gather*}
\operatorname{det}\left(\frac{\partial \Phi_{I^{\prime}}}{\partial u_{j^{\prime \prime}}}\right) \neq 0 ;  \tag{25}\\
\operatorname{det}\left(\begin{array}{ll}
\frac{\partial \Phi_{I^{\prime}}}{\partial u_{j^{\prime \prime}}} & \frac{\partial \Phi_{I^{\prime \prime}}}{\partial u_{j^{\prime \prime \prime}}} \\
\frac{\partial \Phi_{j^{\prime \prime \prime}}}{\partial u_{j^{\prime \prime}}} & \frac{\partial \Phi_{I^{\prime \prime}}}{\partial u_{j^{\prime \prime \prime}}}
\end{array}\right) \neq 0 .
\end{gather*}
$$

From (25) $)_{3}$ we see that we can take $q_{1}=u_{1}, \ldots, q_{N-M}=u_{N-M}$ and the system (17) in the independent variables $u_{1}, \ldots, u_{N-M}, \psi_{1}, \ldots, \psi_{n-(N-M)}$ is hyperbolic. Let us consider the invertible change of variables, expressing the above ones in terms of $u_{1}, \ldots, u_{N-M}, u_{N-M+1}, \ldots, u_{n}$, by means of the transformation law

$$
\begin{equation*}
\Phi_{I^{\prime}}(\underline{u})=\psi_{I^{\prime}} \text { for } I^{\prime}=1,2, \ldots, n-(N-M) . \tag{26}
\end{equation*}
$$

In other words, we are obtaining $u_{N-M+1}, \ldots, u_{n}$ as functions of the previous variables, from eq. (26).
Obviously, the system (17), expressed in terms of the new variables, is still hyperbolic. By using eq. (26) it becomes

$$
\begin{equation*}
\sum_{i=1}^{n} \sum_{j=1}^{N} Z_{k i} A_{i j}^{\alpha} \partial_{\alpha} u_{j}+\sum_{R=1}^{n-(N-M)} \omega^{\alpha} T_{k}^{R} \partial_{\alpha} \Phi_{R}(\underline{u})=\sum_{i=1}^{n} Z_{k i} f_{i} . \tag{27}
\end{equation*}
$$

This system in the independent variables $u_{1}, \ldots, u_{n}$, is hyperbolic. The variables $u_{n+1}, \ldots, u_{N}$ can be obtained from $\Phi_{I^{\prime \prime}}(\underline{u})=0$ as it can be seen from $(25)_{2}$. Consequently, the system (27) is hyperbolic in the variables $u_{1}, \ldots, u_{n}$ constrained only by

$$
\begin{equation*}
\Phi_{I^{\prime \prime}}(\underline{u})=0 \quad \text { for } \quad I^{\prime \prime}=n-(N-M)+1, \ldots, M \tag{28}
\end{equation*}
$$

If we impose that $\Phi_{I}, \underline{(\underline{u})}=0$ (for $I^{\prime}=1, \ldots, n-(N-M)$ ) holds on the initial manifold, it will be satisfied also off it.
The new system (27) has the advantage to have all evolutive equations, even if the variables $\underline{u}$ are still constrained. If we want to eliminate also this drawback, we can proceed in the following way:
Let us consider the system (1) and $N-n$ other scalar-valued equations of the type $\sum_{j=1}^{N} 0_{j}^{\alpha} \partial_{\alpha} u_{j}=0$. We obtain a new system which finally leads to the case $a$. Let us see now some examples of physical applications.

1) Relativistic magnetofluiddynamics.

We have already obtained the system (22) to describe this physical problem; in this system $\psi$ is an auxiliary quantity, and the variables are constrained by $u_{\alpha} b^{\alpha}=0$. We see now, by applying the methodology in case $a$, that the system (22) can be substituted by

$$
\left\{\begin{array}{l}
\partial_{\alpha}\left(n u^{\alpha}\right)=0  \tag{29}\\
\partial_{\alpha}\left[\left(e+p+b^{2}\right) u^{\alpha} u^{\lambda}+\left(p+b^{2} / 2\right) g^{\alpha \lambda}-b^{\alpha} b^{\lambda}\right]+ \\
\quad+u^{\gamma} b^{\lambda} \partial_{\alpha}\left(u^{\alpha} b_{\gamma}-b^{\alpha} u_{\gamma}\right)=0 \\
h^{\beta \lambda} \partial_{\alpha}\left(u^{\alpha} b_{\beta}-b^{\alpha} u_{\beta}\right)+u^{\lambda} u^{\alpha} \partial_{\alpha}\left(u_{\gamma} b^{\gamma}\right)=0
\end{array}\right.
$$

in the variables $e, V^{\alpha}, b^{\alpha}$ which are now not constrained.
2) Relativistic fluiddynamics.

This physical problem is described by the equations

$$
\left\{\begin{array}{l}
\partial_{\alpha}\left(n u^{\alpha}\right)=0  \tag{30}\\
\partial_{\alpha}\left[(e+p) u^{\alpha} u^{\lambda}+p g^{\alpha \lambda}\right]=0
\end{array}\right.
$$

which can be obtained by substituting $b^{\alpha}=0$ in eqs. (21).
If we consider $e, V^{\alpha}$ as variables, this system is not constrained; moreover, it is hyperbolic according to Definition 1. However, to obtain an example of application of the present methodology, we may take $e, n, u^{\alpha}$ as variables constrained by

$$
\begin{equation*}
u^{\alpha} u_{\alpha}+1=0 \tag{31}
\end{equation*}
$$

In this way the number of equations is $n=5$; moreover, $N=6, M=1$. Let $t^{\alpha}=u^{\alpha}$. The matrix $X_{k i}$ is identically null; consequently, $Y_{k i}$ and $Z_{k i}$ are the identity matrix.
Therefore, we are in the case $b$; the system corresponding to (27) is exactly the same initial system (30). However, if we consider the system (30). However, if we cosider the system (30) and the equation $0^{\alpha} \partial_{\alpha} e=0$, we have $n=6, N=6$, $M=1$. The matrices $X_{k i}, Y_{k i}, Z_{k i}, T_{k}^{1}$ are respectively

$$
X_{k i}=\left(\begin{array}{ccc}
0 & 0^{\lambda} & 0 \\
0^{\mu} & 0^{\mu \lambda} & 0 \\
0 & 0^{\lambda} & 1
\end{array}\right) ; Y_{k i}=\left(\begin{array}{ccc}
1 & 0^{\lambda} & 0 \\
0^{\mu} & g^{\mu \lambda} & 0 \\
0 & 0^{\lambda} & 0
\end{array}\right)=Z_{k i} ; T_{k}^{1}=\left(\begin{array}{c}
0 \\
0^{\lambda} \\
1
\end{array}\right)
$$

Therefore the system (17) is given by eqs. (30) and $u^{\alpha} \partial_{\alpha} \psi_{1}=0$, where $\omega^{\alpha}=u^{\alpha}$ has been taken. Finally, the system (24) is given by

$$
\left\{\begin{array}{l}
\partial_{\alpha}\left(n u^{\alpha}\right)=0  \tag{32}\\
\partial_{\alpha}\left[(e+p) u^{\alpha} u^{\lambda}+p g^{\alpha \lambda}\right]=0 \\
u^{\alpha} \partial_{\alpha}\left(u_{\gamma} u^{\gamma}\right)=0
\end{array}\right.
$$

in the variables $e, n, u^{\alpha}$ which are not constrained.
3) Covariant Maxwell electrodynamics.

The field equations are (see e.g. [10])

$$
\begin{equation*}
\partial_{\alpha} F^{\alpha \beta}=-j^{\beta} ; \partial_{\alpha} F^{\alpha \beta^{*}}=0 \tag{33}
\end{equation*}
$$

where

$$
\begin{aligned}
F^{\alpha \beta} & =t^{\alpha} E^{\beta}-t^{\beta} E^{\alpha}+\eta^{\alpha \beta \gamma \delta} H_{\gamma} t_{\delta} \\
F^{\alpha \beta^{*}} & =t^{\alpha} H^{\beta}-t^{\beta} H^{\alpha}-\eta^{\alpha \beta \gamma \delta} E_{\gamma} t_{\delta}
\end{aligned}
$$

The variables are $E^{\alpha}$ (relative electric field) and $H^{\alpha}$ (relative magnetic field), constrained by

$$
\begin{equation*}
\Phi_{1}=t^{\alpha} E_{\alpha}=0 ; \quad \Phi_{2}=t^{\alpha} H_{\alpha}=0 \tag{34}
\end{equation*}
$$

Moreover $t^{\alpha}$ is a field-independent time-direction, such that $t^{\alpha} t_{\alpha}=-1$, and $\eta^{\alpha \beta \gamma \delta}$ is the four-dimensional Levi-Civita's symbol. Therefore we have $n=8, N=8, M=2$,

$$
U_{j j^{\prime}}=\left(\begin{array}{cc}
h_{\gamma}^{\mu} & 0 \\
0 & h_{\gamma}^{\mu}
\end{array}\right), X_{i^{\prime} i}=\left(\begin{array}{cc}
t^{\gamma} & 0^{\gamma} \\
0^{* \gamma} & 0^{* \gamma} \\
0^{\gamma} & t^{\gamma}
\end{array}\right)
$$

$$
Y_{k i}=\left(\begin{array}{cc}
h_{\lambda}^{\nu} & 0 \\
0 & h_{\lambda}^{\nu}
\end{array}\right)=Z_{k i}, T_{k}^{1}=\binom{-t^{\lambda}}{0^{\lambda}}, T_{k}^{2}=\binom{0^{\lambda}}{-t^{\lambda}},
$$

where $0^{* \gamma}$ denotes a $6 \times 4$ matrix with null elements.
Consequently the system (17), in this case, is

$$
\left\{\begin{array}{l}
t^{\alpha} h^{\nu \lambda} \partial_{\alpha} E_{\gamma}+\eta^{\alpha \lambda \gamma \delta} t_{\delta} \partial_{\alpha} H_{\gamma}-t^{\lambda} t^{\alpha} \partial_{\alpha} \psi_{1}=-h^{\lambda \gamma} j_{\gamma} \\
-\eta^{\alpha \lambda \gamma \delta} t_{\delta} \partial_{\alpha} E_{\gamma}+t^{\alpha} h^{\gamma \lambda} \partial_{\alpha} H_{\gamma}-t^{\lambda} t^{\alpha} \partial_{\alpha} \psi_{2}=0 .
\end{array}\right.
$$

The system (24) is

$$
\left\{\begin{array}{l}
t^{\alpha} h^{\gamma \lambda} \partial_{\alpha} E_{\gamma}+\eta^{\alpha \lambda \gamma \delta} t_{\delta} \partial_{\alpha} H_{\gamma}-t^{\lambda} t^{\alpha} \partial_{\alpha}\left(t_{\nu} E^{\gamma}\right)=-h^{\lambda \gamma} j_{\gamma}  \tag{35}\\
-\eta^{\alpha \lambda \gamma \delta} t_{\delta} \partial_{\alpha} E_{\gamma}+t^{\alpha} h^{\gamma \lambda} \partial_{\alpha} H_{\gamma}-t^{\lambda} t^{\alpha} \partial_{\alpha}\left(t_{\gamma} H^{\gamma}\right)=0 .
\end{array}\right.
$$

in the variables $E^{\gamma}, H^{\gamma}$ which are not constrained.
We notice that (35) can be written also as

$$
\begin{equation*}
\partial_{\alpha}\left(F^{\alpha \lambda}+E^{\alpha} t^{\lambda}\right)=-h^{\lambda \gamma} j_{\gamma} ; \quad \partial_{\alpha}\left(F^{\alpha \lambda^{*}}+H^{\alpha} t^{\lambda}\right)=0, \tag{36}
\end{equation*}
$$

where also the conservative form is preserved.

## 5. Hyperbolicity in every time-like direction.

The problem of characterizing the system, which are hyperbolic in every time-like direction, is still open. Strumia has obtained a very interesting result, in the case of symmetric hyperbolic systems (see appendix of ref. [11]). The general case remains still to be investigated. Here I rest content of the following result. Let us choose a particular $t_{\alpha}$ and find the matrix $Z_{k i}$ depending on this $t_{\alpha}$. Then the condition 2) of defintion 2 holds for every other time-direction $t_{\alpha}^{*}$ if the following two condition are satisfied

1) the system (10) is hyperbolic in the time-direction $t_{\alpha}$;
2) the characteristic velocities of system (10) (i.e., the solutions $\lambda$ of condition 3) in Definition 2) do not exceed the speed of light.

In fact, let us assume, by absurd, that a four-vector $t_{\alpha}^{*}$ exists, such that $t_{\alpha}^{*} t^{* \alpha}=$ -1 and the system

$$
t_{\alpha}^{*} \sum_{i=1}^{n} \sum_{j=1}^{N} Z_{k i} A_{i j}^{\alpha} \delta u_{j}=0, \quad \sum_{j=1}^{N} \frac{\partial \Phi_{I}}{\partial u_{j}} \delta u_{j}=0,
$$

in the independent unknowns $\delta u_{j}$, has a non null solution $\delta u_{j}^{*}$.

Let $n^{\alpha}$ be defined by $t_{\alpha}^{*}=-t^{\gamma} t_{\gamma}^{*} t_{\alpha}+n_{\alpha}$.
We have $t_{\alpha} n^{\alpha}=0$ and $n_{\alpha} n^{\alpha} \geq 0$, if $n_{\alpha} n^{\alpha}=0$ it would follow that $n^{\alpha}=0$, and condition 2) of Definition 2 would be violated; consequently we have $n_{\alpha} n^{\alpha}>0$. Let us define $\zeta_{\alpha}=n_{\alpha}\left(n_{\gamma} n^{\gamma}\right)^{-1 / 2}$.
From condition 2) we have that the eigenvalues $\lambda$ of the problem

$$
\left(\zeta_{\alpha}-\lambda t_{\alpha}\right) \sum_{i=1}^{n} \sum_{j=1}^{N} Z_{k i} A_{i j}^{\alpha} \delta\left(u_{j}=0, \sum_{j=1}^{N} \frac{\partial \Phi_{I}}{\partial u_{j}} \delta u_{j}=0\right.
$$

are such that $|\lambda| \leq 1$.
Now this problem is equivalent to

$$
\left[n_{\alpha}-\lambda t_{\alpha}\left(n_{\gamma} n^{\gamma}\right)^{1 / 2}\right] \sum_{i=1}^{n} \sum_{j=1}^{N} Z_{k i} A_{i j}^{\alpha} \delta u_{j}=0, \sum_{j=1}^{N} \frac{\partial \Phi_{I}}{\partial u_{j}} \delta u_{j}=0
$$

which has the solution $\lambda=t^{\gamma} t_{\gamma}^{*}\left(n_{\delta} n^{\delta}\right)^{-1 / 2}, \delta u_{j}=\delta u_{j}^{*}$.
(See what we assumed by absurd). Therefore we have
$\left|t^{\gamma} t_{\gamma}^{*}\right|\left(n_{\delta} n^{\delta}\right)^{-1 / 2} \leq 1$, from which $\left(t^{\gamma} t_{\gamma}^{*}\right)^{2} \leq-1+\left(t^{\delta} t_{\gamma}^{*}\right)^{2}!!$
This absurd result proves our statement.
I conclude this section noticing that the characteristic velocities of system (17) are those of system (10) and $\lambda^{\prime}=\left(n^{\alpha} \omega_{\alpha}\right) /\left(t^{\alpha} \omega_{\alpha}\right)$. (See in section III). The first ones, of these, do not exceed the speed of light. Regarding $\lambda^{\prime}$, we have that $\left(\lambda^{\prime}\right)^{2} \leq 1$ holds iff $\left(n^{\alpha} \omega_{\alpha}\right)^{2} \leq\left(t^{\alpha} \omega_{\alpha}\right)^{2}$ for every $n^{\alpha}$ such that $n^{\alpha} n_{\alpha}=1, n^{\alpha} t_{\alpha}=0$. This relation can be written in the references frames where $t_{\alpha}, \omega_{\alpha}$ have the components $t_{\alpha} \equiv(1,0,0,0), \omega_{\alpha} \equiv\left(\omega_{0}, \omega_{1}, 0,0\right)$; it reads $\left(n_{1} \omega^{1}\right)^{2} \leq\left(\omega^{0}\right)^{2}$ for every $n_{i}$ such that $n_{i} n^{i}=1$. Now $\left(n_{1} \omega^{1}\right)^{2}$ assumes its maximum value for $n_{1}=1, n_{2}=n_{3}=0$; this maximum is $\left(\omega^{1}\right)^{2}$. Therefore we must have $\left(\omega^{1}\right)^{2} \leq\left(\omega^{0}\right)^{2}$, i.e., $\omega^{\alpha} \omega_{\alpha} \leq 0$.
Consequently $\omega^{\alpha}$ must be chosen as a time-like or a light-like 4 -vector. If we choose $\omega^{\alpha}=t^{\alpha}$, this condition is surely satisfied.

## Appendix 1.

I prove here that the matrix $t_{\alpha} \sum_{j=1}^{N} A_{i j}^{\alpha} U_{j j^{\prime}}$, has rank $N-M$. Let $\delta Q_{j}$ be a solution of the system

$$
\begin{equation*}
t_{\alpha} \sum_{j^{\prime}, j=1}^{N} A_{i j}^{\alpha} U_{j j^{\prime}} \delta Q_{j^{\prime}}=0 \tag{A.1}
\end{equation*}
$$

If $\delta P_{j}$ is defined by $\delta P_{j}=\sum_{j^{\prime}=1}^{N} U_{j j^{\prime}} \delta Q_{j^{\prime}}$, we have that $t_{\alpha} \sum_{j=1}^{N} A_{i j}^{\alpha} \delta P_{j}=0$, $\sum_{j=1}^{N} \frac{\partial \Phi_{I}}{\partial u_{j}} \delta P_{j}=0$, from which $t_{\alpha} \sum_{i=1}^{n} \sum_{j=1}^{N} Z_{k i} A_{i j}^{\alpha} \delta P_{j}=0, \sum_{j=1}^{N} \frac{\partial \Phi_{I}}{\partial u_{j}} \delta P_{j}=0 ;$ from condition 2) of Definition 2, it follows that $\delta P_{j}=0$, i. e.,

$$
\begin{equation*}
\sum_{j^{\prime}=1}^{N} U_{j j^{\prime}} \delta Q_{j^{\prime}}=0 \tag{A.2}
\end{equation*}
$$

Vice versa if $\delta Q_{j}$, is a solution of system (A.2), then it satisfies also (A.1). Therefore the systems (A.1) and (A.2) have the same solutions; to this end it is necessary that the matrices of their coefficients have the same rank $\rho$ (We have $N-\rho$ free unknowns). Consequently, $t_{\alpha} \sum_{j=1}^{N} A_{i j}^{\alpha} U_{j j^{\prime}}$ and $U_{j j^{\prime}}$ have the same rank, i.e., $N-M$.

## Appendix 2.

I prove now that the $2 n \times n$ matrix $\binom{Y_{i^{\prime} i}}{X_{i^{\prime} i}}$ has rank $n$. Let $C_{h i}$ be $N-M$ 1.i. solutions of $\sum_{i=1}^{n} X_{i^{\prime} i} C_{h i}=0$; these solutions can be chosen also orthonormal. From system (9) it follows that

$$
\begin{equation*}
Y_{k i}=\sum_{h=1}^{N-M} \lambda_{k h} C_{h i} \tag{A.3}
\end{equation*}
$$

If $\lambda_{k 1}$ is a linear combination of $\lambda_{k 2}, \ldots, \lambda_{k(N-M)}$, i.e., $\lambda_{k 1}=\sum_{p=2}^{N-M} q_{p} \lambda_{k p}$, then (A.3) gives $Y_{k i}=\sum_{p=2}^{N-M} \lambda_{k p}\left(C_{p i}+q_{p} C_{1 i}\right)$, i.e., $Y_{k i}$ are linear combinations of $N-M-1$ vectors; this is not possible because $Y_{k i}$ has rank $N-M$. Similarly no other $\lambda_{k h}$ is a linear combination of the remaining ones. In other words, the matrix $L$ with elements $\lambda_{k h}$ has rank $N-M$.
Let us consider now the system

$$
\begin{equation*}
\sum_{i=1}^{n} Y_{i^{\prime} i} x_{i}=0, \sum_{i=1}^{n} X_{i^{\prime} i} x_{i}=0 \tag{A.4}
\end{equation*}
$$

and prove that it has only the solution $x_{i}=0$. In this way our aim will be achieved. From (A.4) $)_{2}$ we have $x_{i}=\sum_{h=1}^{N-M} u_{h} C_{h i}$; by substituting this and (A.3) in (A.4) ${ }_{1}$ we obtain

$$
\sum_{h, r=1}^{N-M} \lambda_{i^{\prime} h} C_{h i} \mu_{r} C_{r i}=0, \quad \text { i.e., } \quad L C C^{T} \underline{u}=0
$$

This system can be written as

$$
\begin{equation*}
L \underline{u}=0 \tag{A.5}
\end{equation*}
$$

because $C C^{T}=I_{N-M}$.
Now $L$ is a $n X(N-M)$ matrix with rank $N-M$; therefore we have $\underline{u}=0$ from (A.5), i.e., $x_{i}=0$.

## Appendix 3.

Let us prove the following theorem
Let us consider a $n \times n$ matrix $A$ and $a n \times p$ matrix $B$; moreover let $A B$ have rank $p$. It follows that $A$ has rank $\rho \geq p$.
In fact the system $A^{T} \underline{q}=0$ has $n-\rho 1$ i. solutions $\underline{q}_{1}, \ldots, \underline{q}_{n-\rho}$. Let $Q$ be an invertible $n \times n$ matrix whose first $n-\rho$ rows are $\underline{q}_{1}^{T}, \ldots, \underline{q}_{n-\rho}^{T}$. It follows that $Q A B$ has rank $p$ and its first $n-\rho$ rows have null elements; the remaining rows are $\rho$ in number and, consequently, $p \leq \rho$. By applying this theorem with $A=\sum_{i^{\prime}=1}^{n} \lambda_{k i^{\prime}} Y_{i^{\prime} i}$ and $B=t_{\alpha} \sum_{i=1}^{n} \sum_{j=1}^{N} A_{i j}^{\alpha} \frac{\partial u_{j}}{\partial q_{h}}$, we have that the rank $\rho$ of $\sum_{i^{\prime}=1}^{n} \lambda_{k i^{\prime}} Y_{i^{\prime} i}$ is such that $\rho \geq N-M$. Now the rank of $Y_{i^{\prime} i}$ is $N-M$, from which $\rho \leq N-M$. Finally, we have $\rho=N-M$, a property which I used in section II.

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